

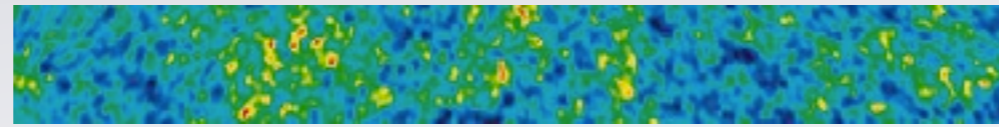
The CFHT Stripe 82 Survey (CS82)



Alexie Leauthaud, Jean-Paul Kneib, Ludovic van Waerbeke,
Martin Makler, and the CS82 collaboration

CMB-Galaxy Lensing Cross-correlations

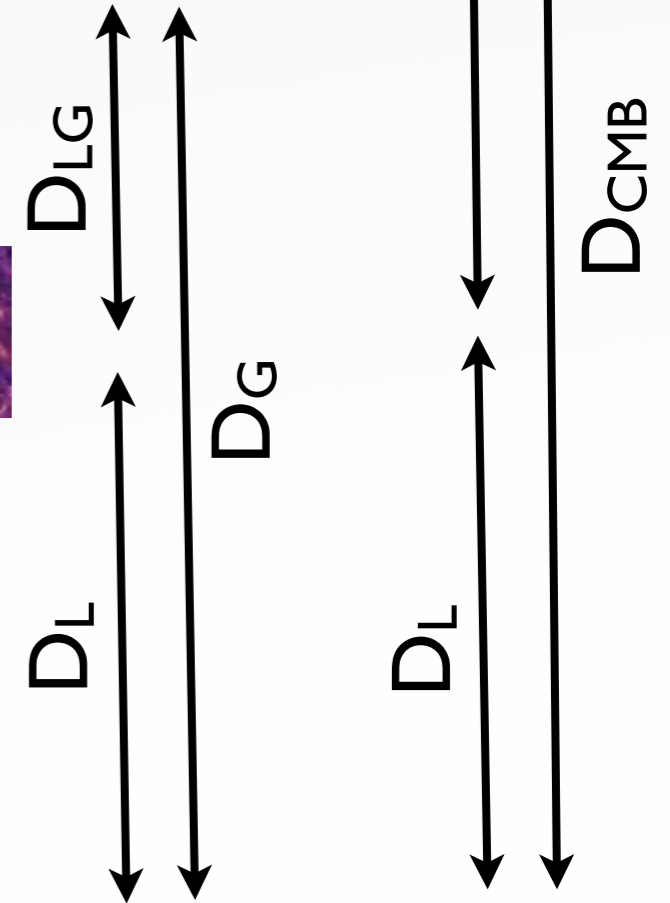
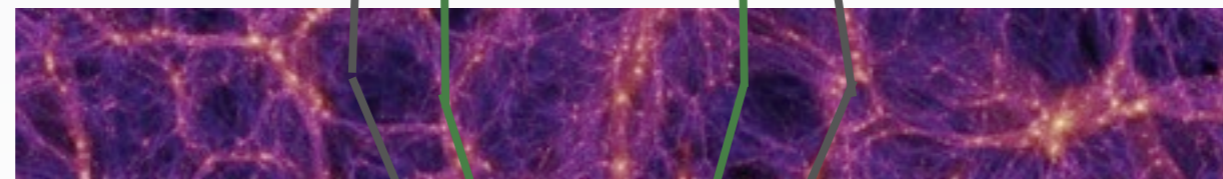
with: Sudeep Das, Eric Linder, Charlotte Welker



$$\kappa = \Sigma / \Sigma_{\text{crit}}$$

$$\Sigma_{\text{crit}} \sim D_S / D_L D_{LS}$$

Σ = surface mass density



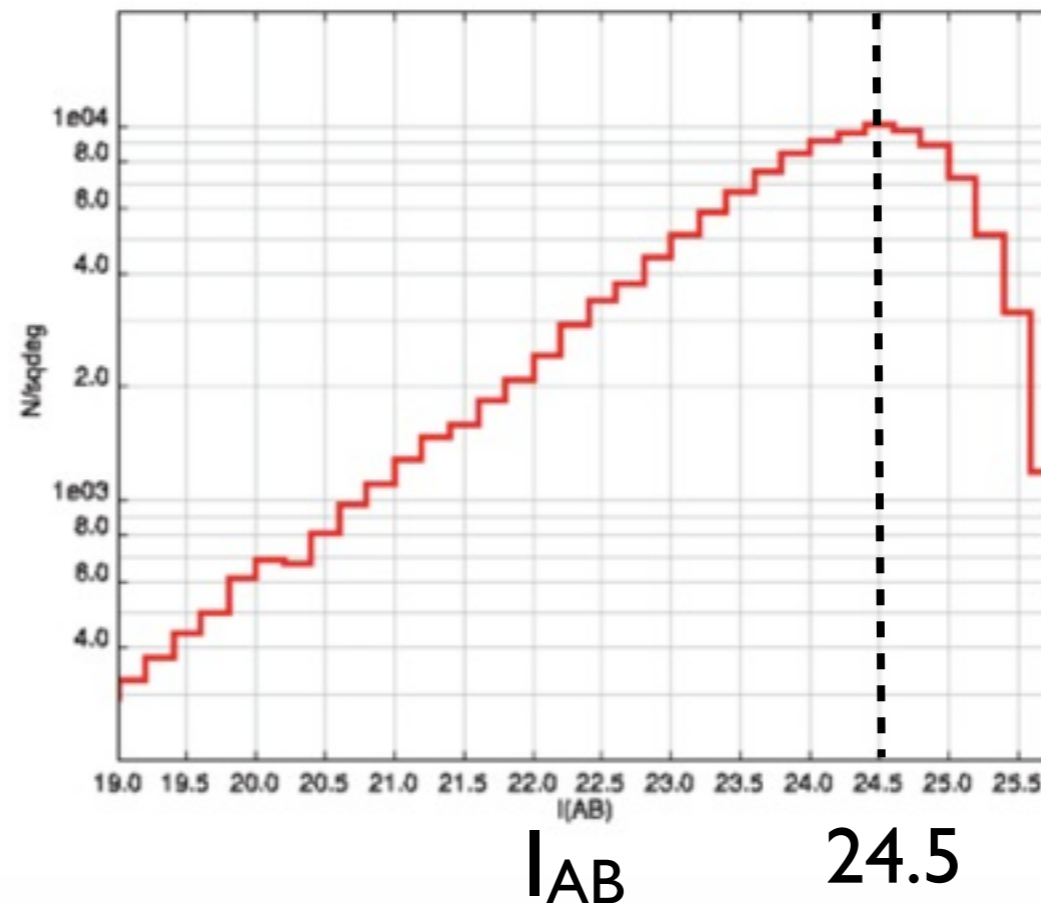
- Amplitude of power spectrum at Z_L
- $R = (D_{L-CMB} D_G) / (D_{LG} D_{CMB})$



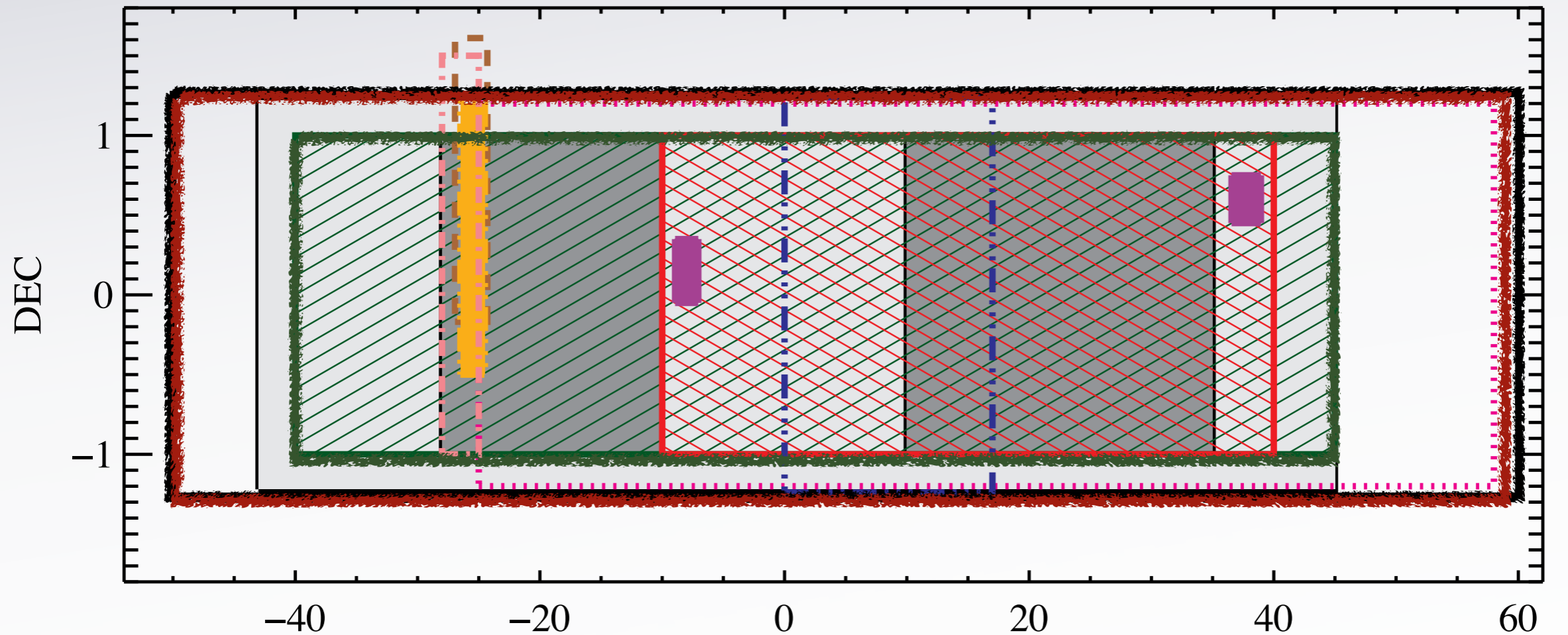
Das et al. 2008

CS82 in a nutshell







- 170 MegaCam (1 sq.degree) pointings with slight overlaps
- Same area as CFHTLS wide field !!
- Mostly between $-40 < \text{Ra} < +45$ and $-1 < \text{Dec} < +1$, “Stripe 82”
- 30 minutes observations, 4x410s science images
- I-band depth: ~ 23.5 mag for galaxies, ~ 24 mag for point sources



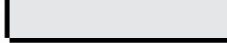



Where is Stripe 82 ?



Imaging :

-  Stripe 82 SDSS imaging ($ugriz, i < 22.75, 270 \text{ deg}^2$)
-  CFHT Stripe-82 survey ($170 \text{ deg}^2, i < 23.5, \text{seeing} < 0.8''$)
-  UKIDSS LAS, $K_{\text{vega}} = 18.4$
-  UKIDSS: DXS Field 4
-  CFHTLS W4
-  WISER (this proposal), $K_{\text{AB}} = 22$

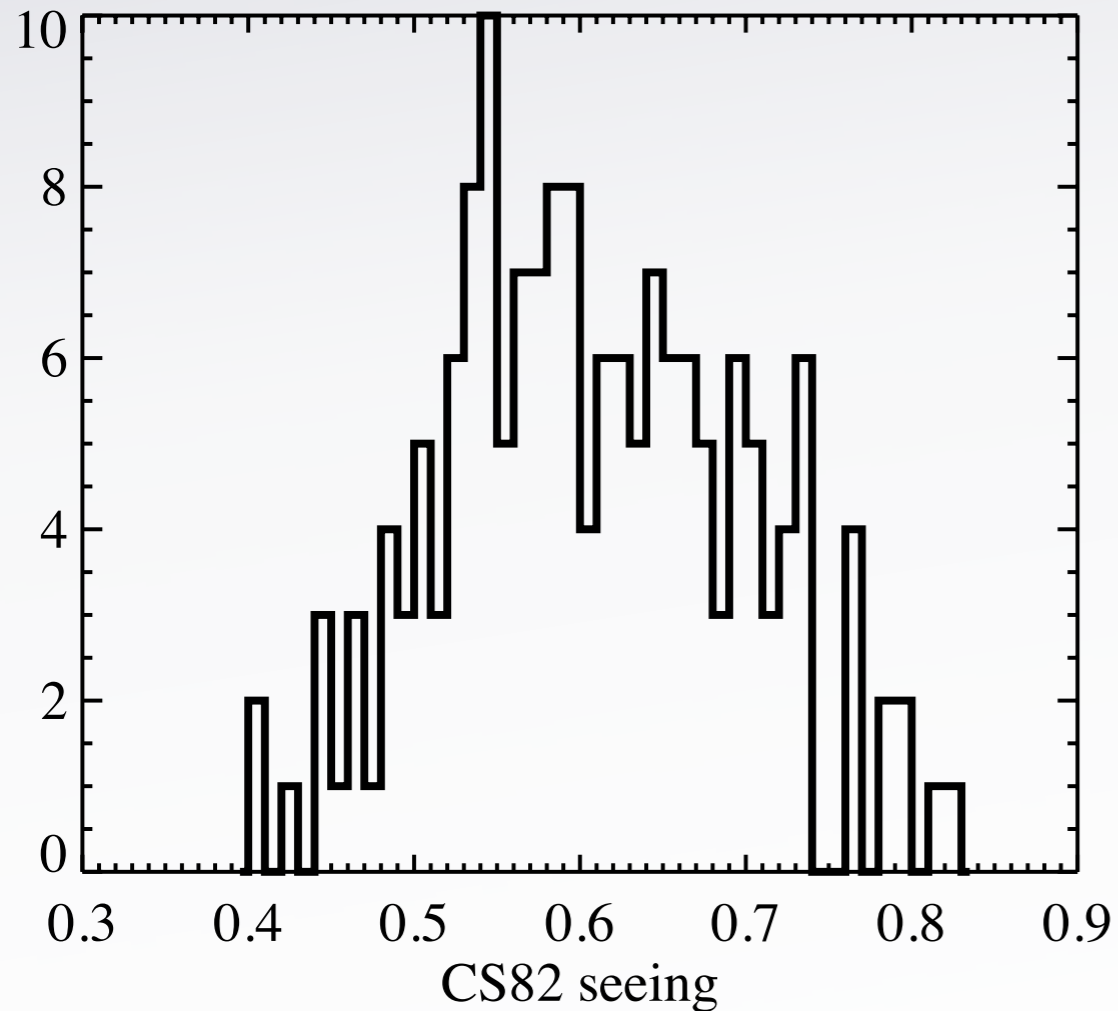
Spectroscopy :

-  BOSS ($220 \text{ deg}^2, 40,000$ redshifts)
-  DEEP2 and PRIMUS
-  VVDS
-  Wiggle-z

Radio :

-  VLA

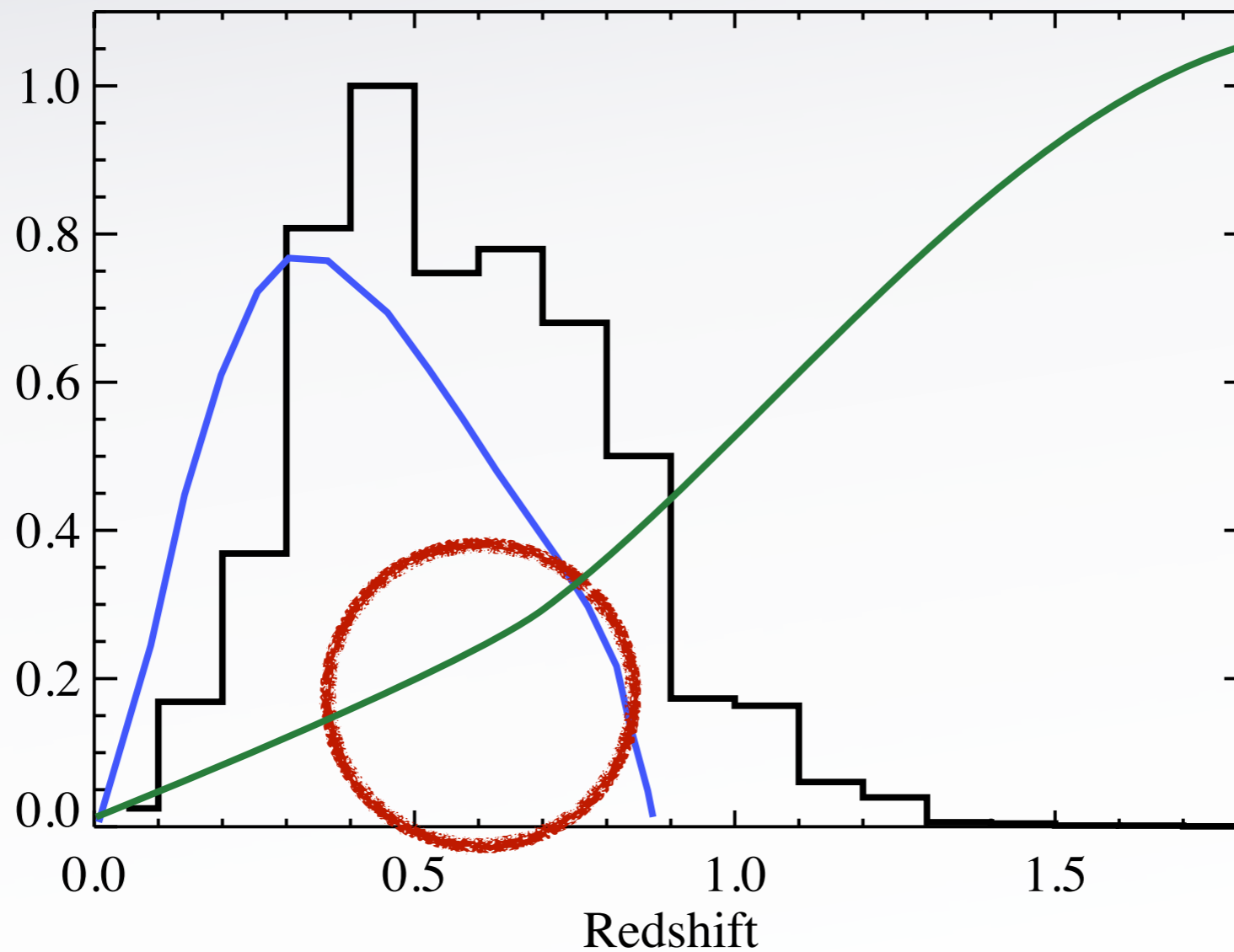
CS82 observing conditions



CS82 mean seeing = 0.6''
Good for kappa maps with galaxy lensing

CS82 Redshift Distribution

CS82 galaxies for lensing

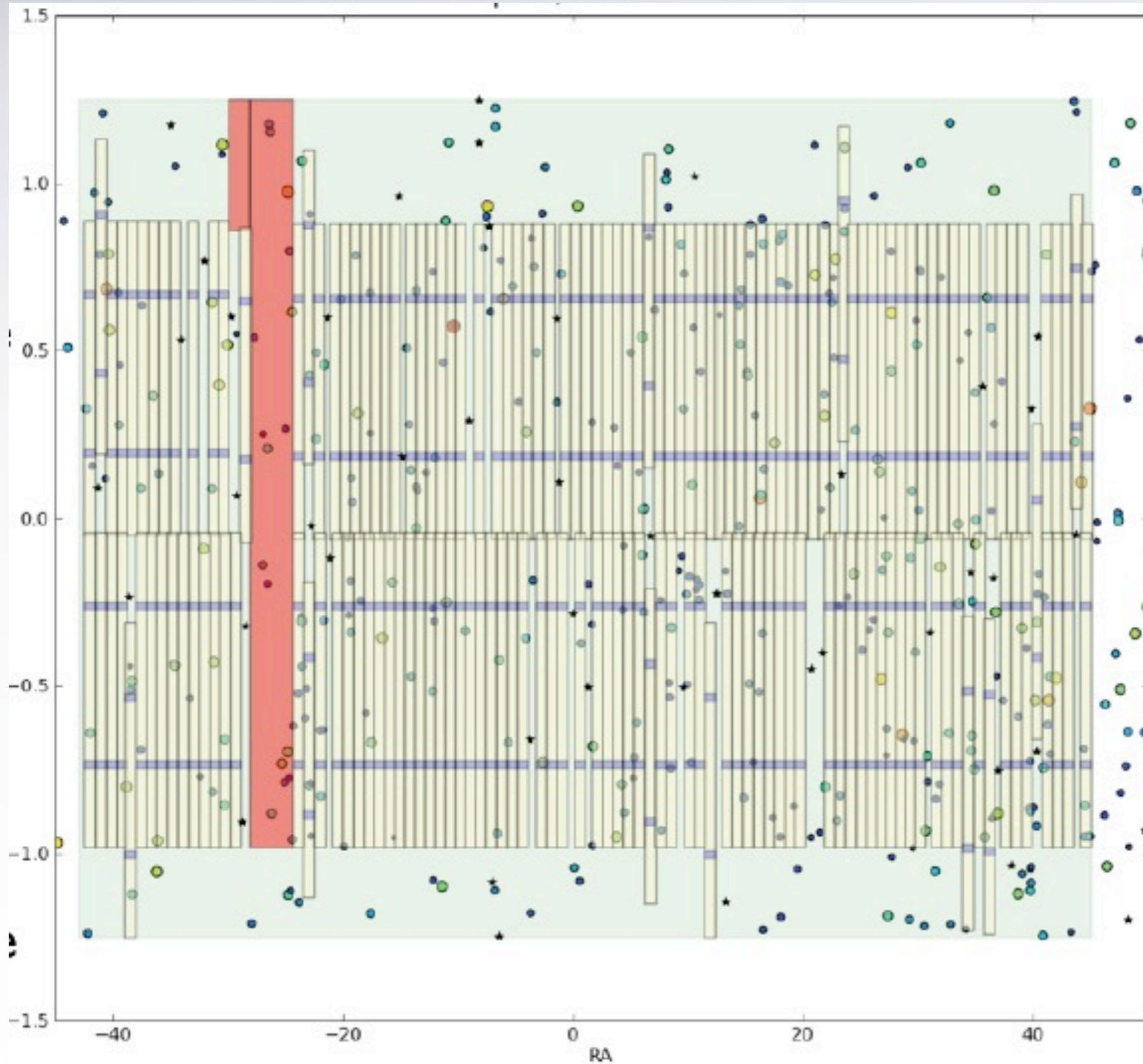


Predicted:
7 sigma detection

Galaxy 'regular' lensing

CMB lensing

Observing strategy



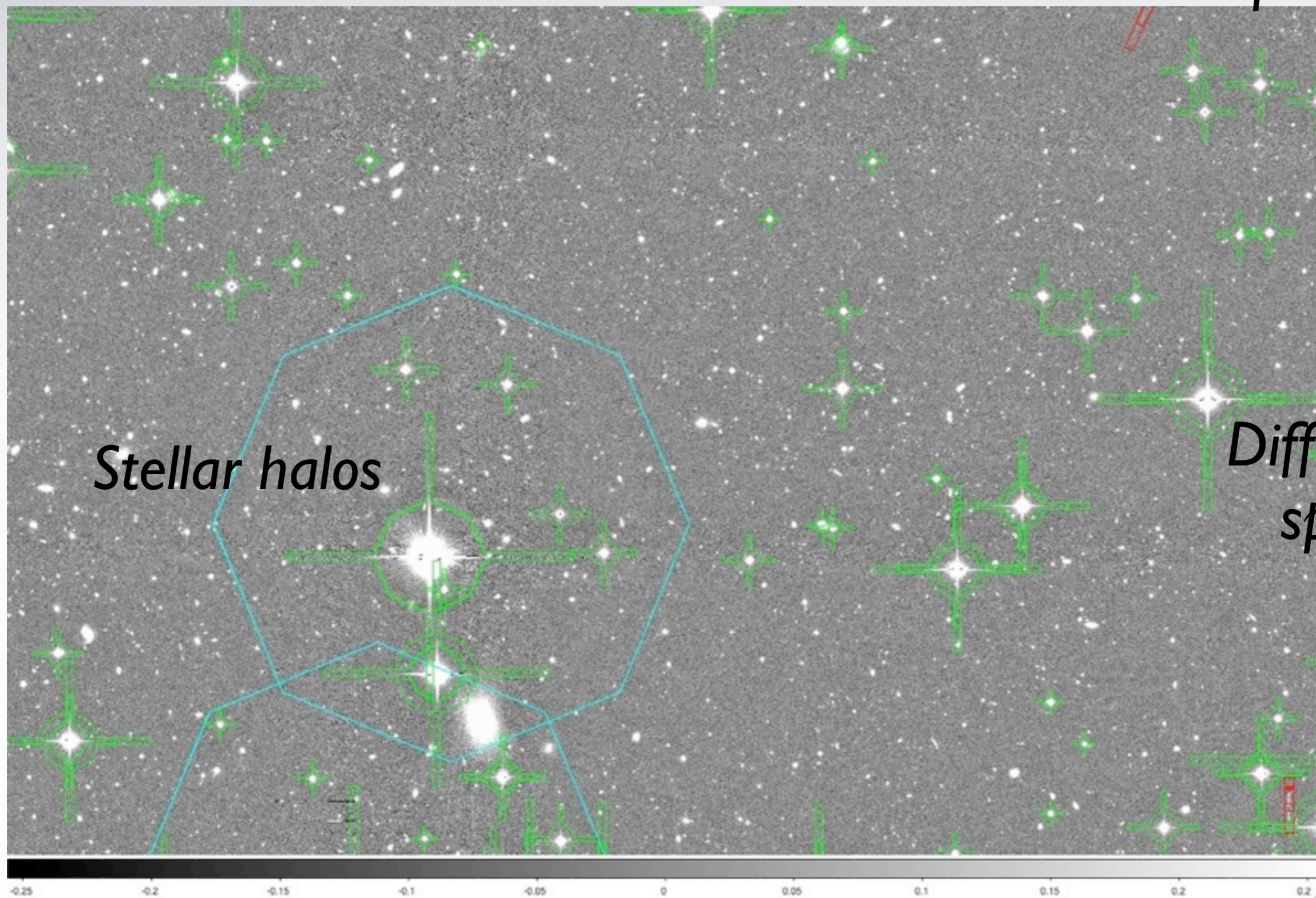
Observing strategy such that we have tried to avoid massive stars

CS82 timeline: fast track

- Proposal accepted spring 2010
- Observations from August 2010 to January 2011, service mode observations insuring good seeing for the survey (seeing < 0.8"), 100% observations complete
- First data reduction complete
- Mask making
- Source galaxy number density: ~ 10 per armin^2

Mask making: done!

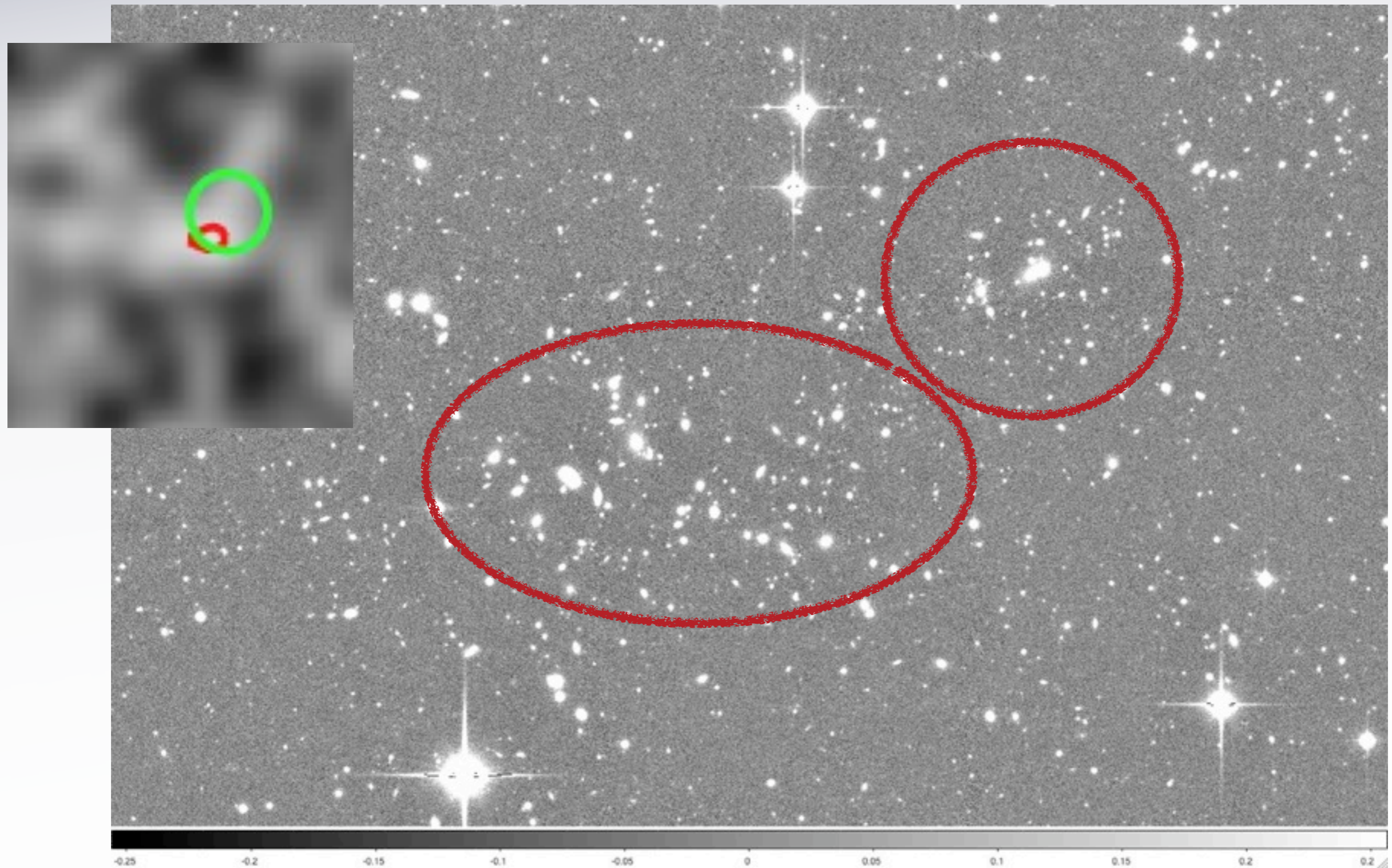
*Asteroids/image
defects*



Stellar halos

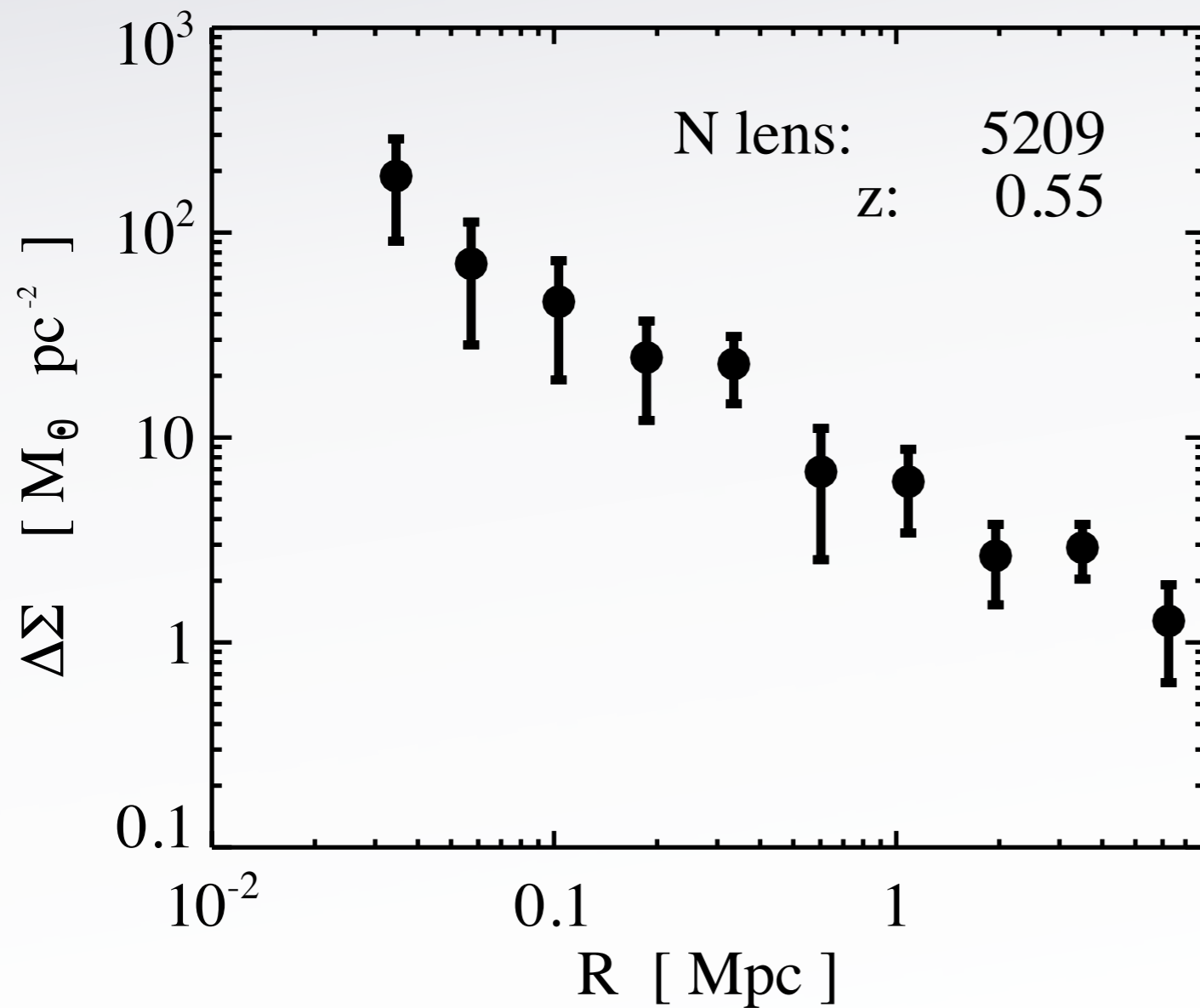
*Diffraction
spikes*

What I will show you next time



$z=0.4$ cluster, $N_{\text{gal}} \sim 15$

Yes, we have strong lensing signals ...



galaxy-galaxy lensing
around a sample of
spectroscopic lens
galaxies at $z=0.5$ from
BOSS

Summary

- CS82: “traditional” weak lensing survey, 170 deg²
- Covers Stripe 82 - overlaps with ACT
- In the process of making kappa maps
- CMB-galaxy lensing cross-correlations
- Stay tuned!

